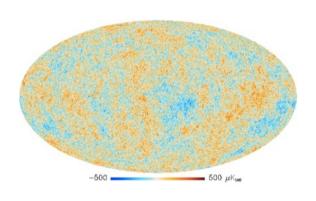
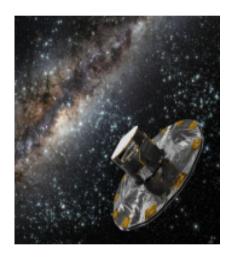
### From the Big Bang to Big Data

### Ofer Lahav (UCL)













### Outline

- What is 'Big Data'?
- What does it mean to computer scientists
  vs physicists?
- The Alan Turing Institute
- Machine learning examples from Astronomy
- The next big projects
- Challenges

# What is 'Big Data'?

- Wikipedia's definition: "data sets that are so large or complex that TRADITIONAL data processing applications are inadequate to deal with them".
- Clearly, this is a 'moving target'.
- "Big data is high volume, high velocity, and/or high variety information assets that require new forms of processing to enable enhanced decision making, insight discovery and process optimization." (Gartner)

## The Alan Turing Institute

- Founding universities:
  Cambridge, Edinburgh, Oxford, UCL, Warwick
- Based at the British Library
- See summary of an ATI summit held in Jan 2016 at the RS on "Big Data in Physical Sciences":

https://indico.cern.ch/event/449964/overview

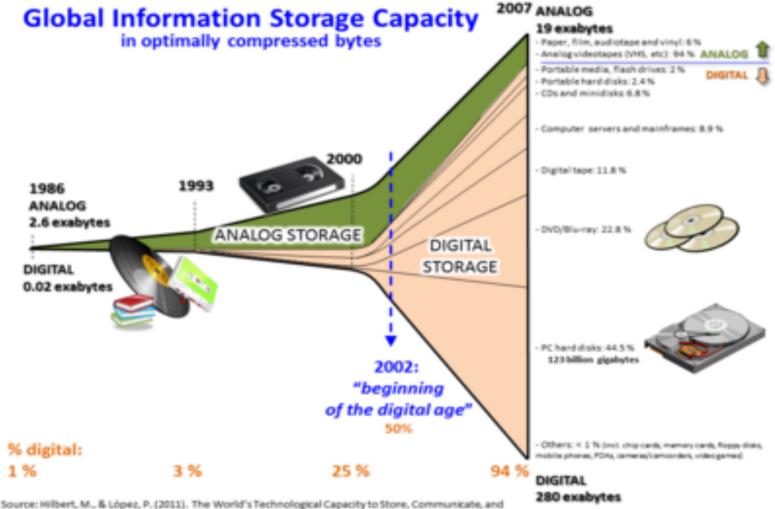
## Astrophysics and HEP examples

- Google: 3.5 Billion Google searches per day
- LHC: 600 Million collisions per second (only 100 per second are 'interesting')
- SDSS: 200 Giga(10^9) Bytes per night
- DES: 1 Tera (10^12) Bytes per night
- LSST: 15 Tera Bytes per night
- SKA: 1 Peta (10^15) Bytes per day

### **Big numbers**

- Exo-planets: 9-1 (+1?) +2000+...
- Gaia: 1B stars
- DES: 300M galaxies
- Euclid/LSST: 1B galaxies
- Simulations: N-body, Hydro (many times the data)

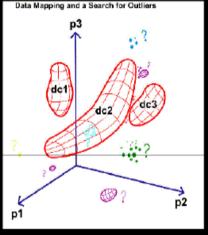
### Information storage



Compute Information. Science, 332(6025), 60 -65. http://www.martinhilbert.net/WorldinfoCapacity.html

### Big Data Science: Scientific KDD (Knowledge Discovery from Data)

- Characterize the known (clustering, unsupervised learning)
- Assign the new (classification, supervised learning)
- Discover the unknown (outlier detection, semi-supervised learning)



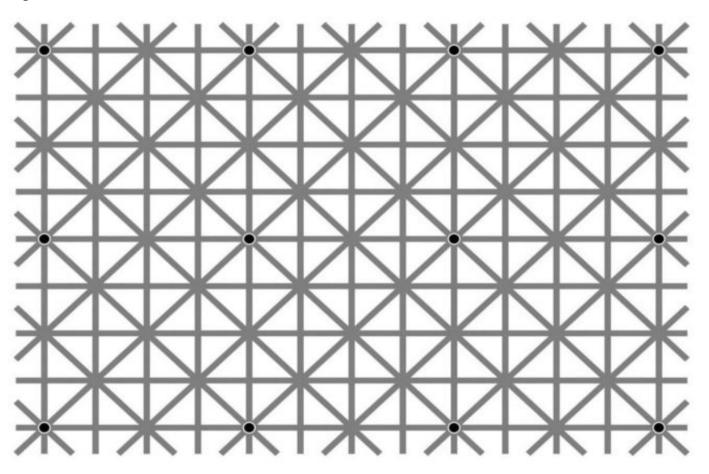
Graphic from S. G. Djorgovski

### Benefits of very large datasets:

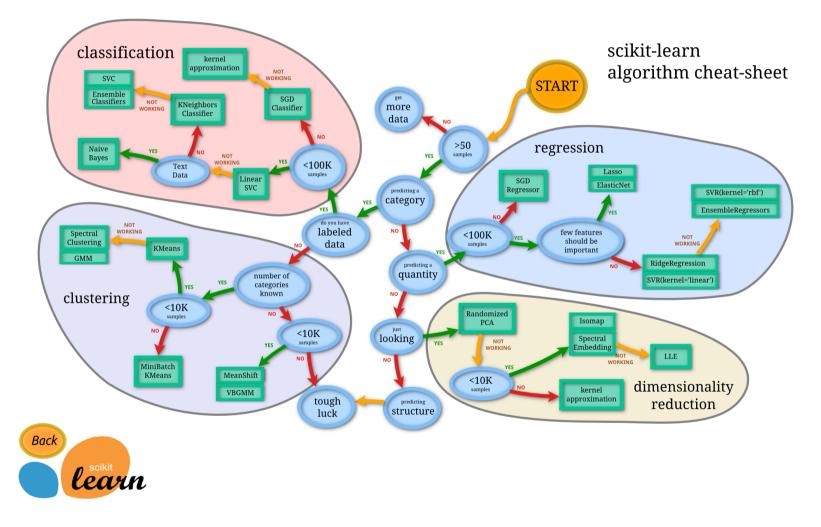
- best statistical analysis of "typical" events
- automated search for "rare" events

# Can we trust just the human brain?

(can you see 12 black dots at once?)



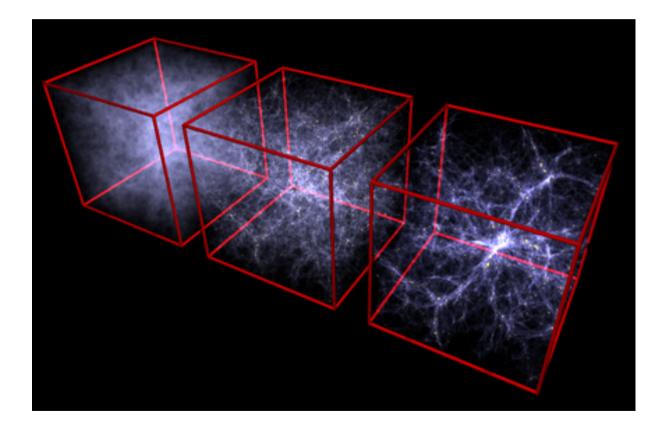
# Can we trust Machine Learning?



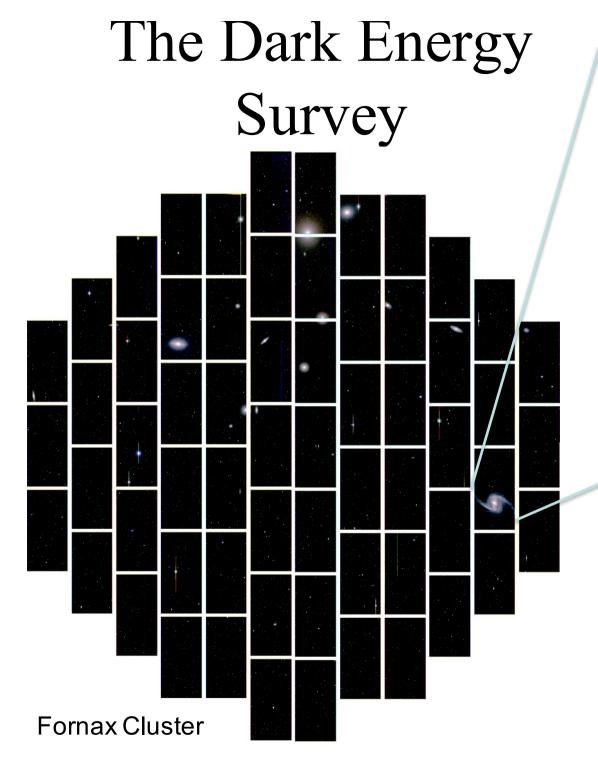
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# Big Data in mock universes: How to contrast with the data?

e.g. intensive Bayesian + MCMC approaches



Springel et al.





NGC 1365

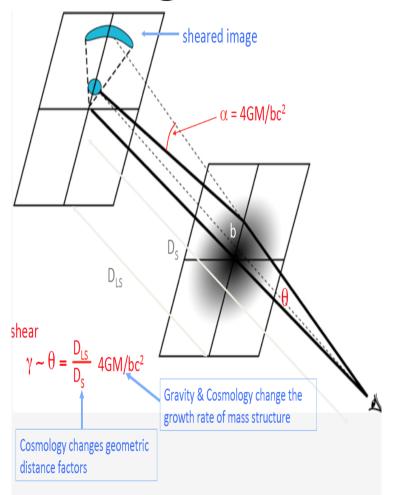
	Objects	As of Dec 2015	Expected from full 5yr DES
.00329	Galaxies with photo-z (> 10 sigma)	7M (SV), 100M (Y1+Y2),	300M
	Galaxies with shapes	3M (SV), 80M (Y1+Y2)	200M
	Galaxy clusters (lambda>5)	150K (Y1+Y2)	380K
	SN la SLSN	1000 2 + confirmed + candidates	Thousands 15-20
	New Milky Way companions	17	25
	QSO's at z> 6 Lensed QSO's	<ul><li>1 + confirmed + candidates</li><li>2 + candidates</li></ul>	375 100 (i<21)
arXIV:1601	Stars (> 10 sigma) Solar System: Trans	2M (SV), 30M (Y1+Y2)	100M
	Jupiter Trojans Main Belt asteroids	32 in SN fields + 2 in the WF 19 300K (Y1+Y2)	50 + many more in the wide field
	Kuiper Belt Objects		500-1000

DES Non—DE Overview arXiv:1601.00329

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### Gravitational Lensing: Weak and Strong

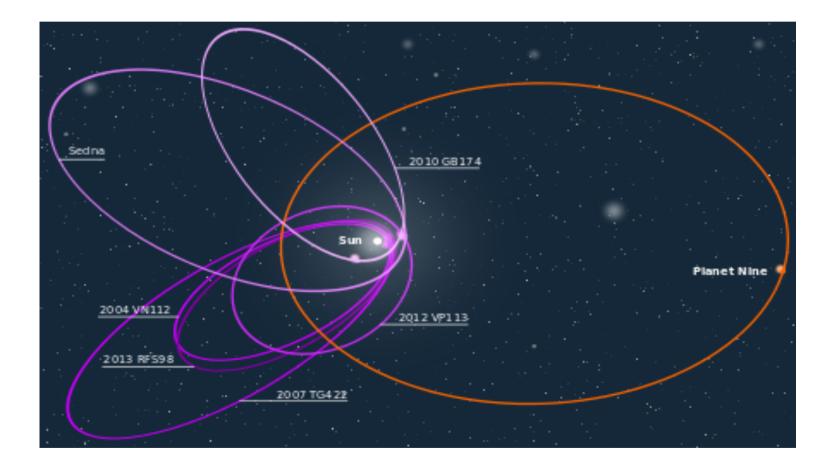




#### HST CLASH cluster MACS1206

#### DES Mass Map from Weak Lensing 1.0% DARK ENERGY SURVEY cluster richness $\lambda$ Galaxy clusters Ο 0 0 $\bigcirc$ 80 160 $-45^{\circ}$ 20 40 00 matter density $\kappa_E$ [compared to cosmic mean] SO 0.5% 30 Mpc $-50^{\circ}$ 00 0 °00° 00 00 0.0% 8 0:0 8 0 0 )0 80 000 -55° 00 6 000 -0.5% 08 0 ood 0 0 00 Ο 0 0 0 õ 0 $-60^{\circ}$ 00 Chang, et al -1.0% 6:30h 5:30h 6:00h 5:00h 4:00h Vikram, et al

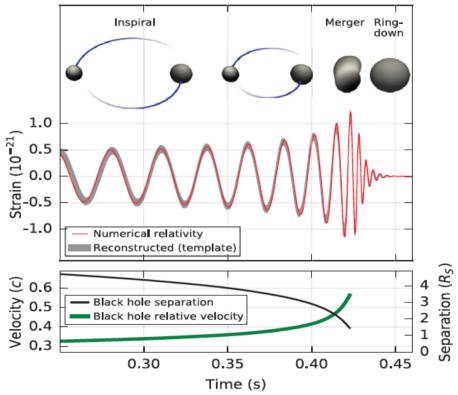
### The search for Planet 9 (one of the 6 minor planets discovered by DES)



David Gerdes et al, DES TMO WG

# The new window of Gravitational Waves:

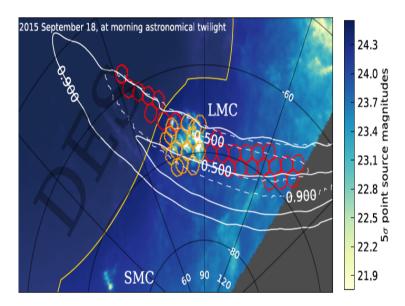
### 3 events detected so far; on per month expected; and from 2019 a few per week



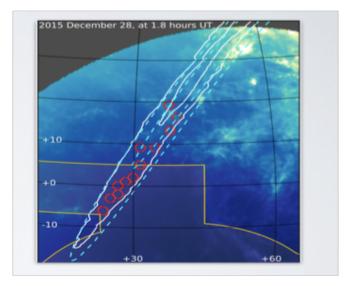
LIGO collaboration 2016

### DES LIGO GW follow ups

### GW150914

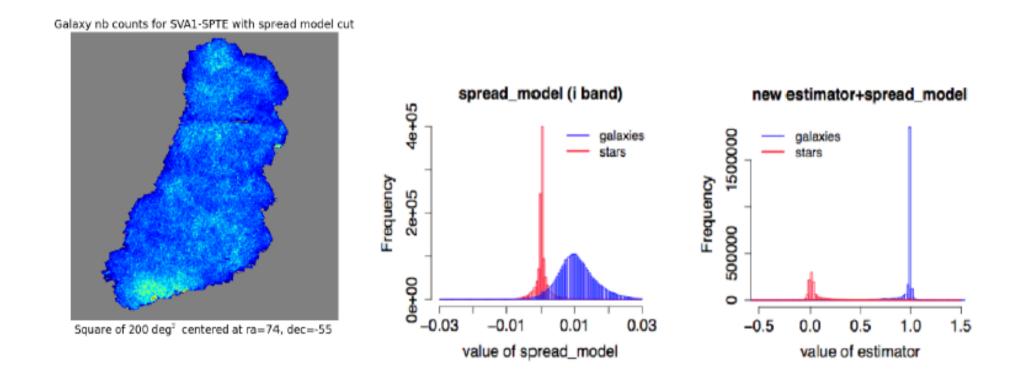


#### GW151226



Soares-Santos et al. (2016) Annis et al. (2016) Abbott et al. (2016) Cowperthwaite et al. (2016)

### Star/galaxy separation in DES



Soumagnacet al (1306.5236)

# GALAXY ZOO

• One Million galaxies classified by 100,000 people! Is the galaxy simply smooth and rounded, with no sign of a disk?



Smooth



Features or disk

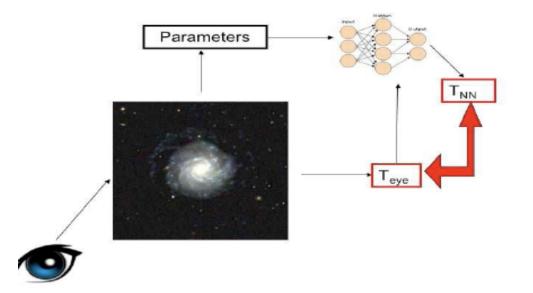


Star or artifact



Lintott et al.

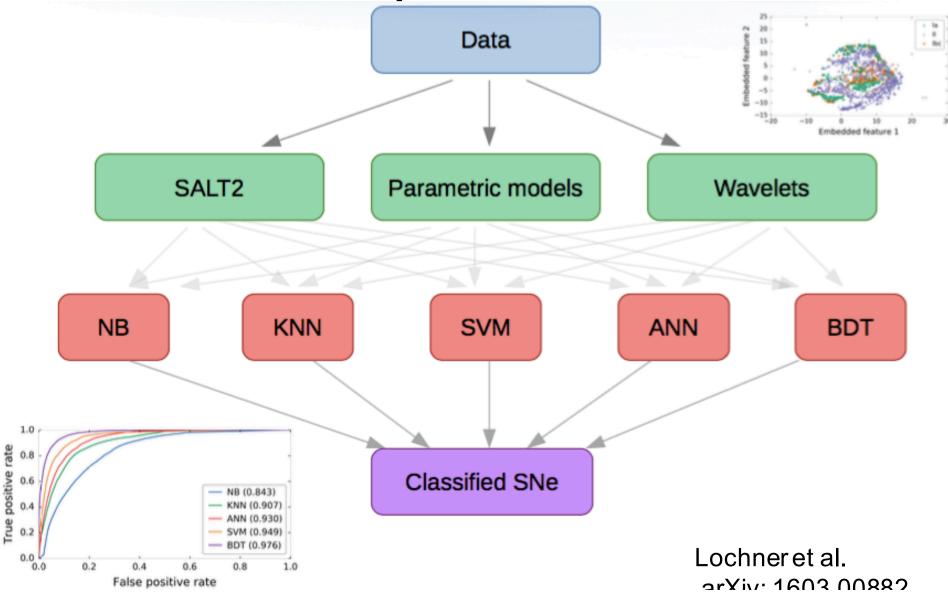
### Galaxy zoo and machine learning



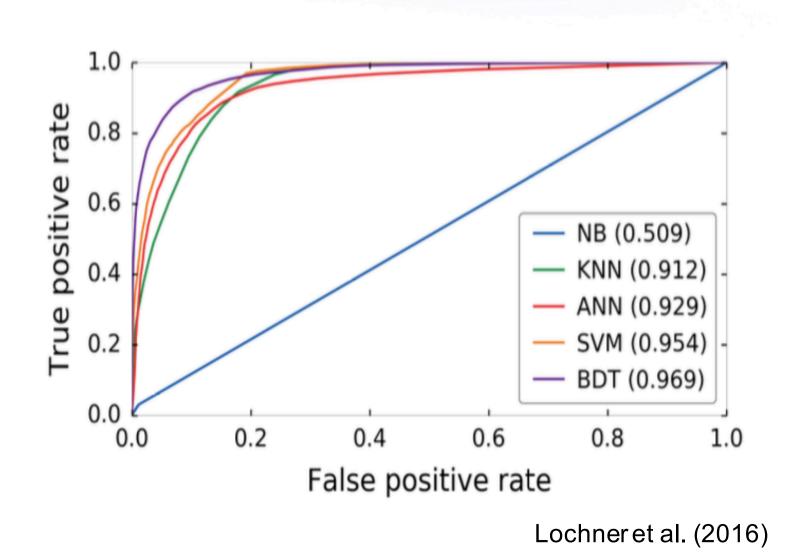
		GALAXY ZOO		
		Elliptical	Spiral	Star/Other
Α	ELLIPTICAL	91%	0.08%	0.5%
Ν	SPIRAL	0.1%	93%	0.2%
Ν	STAR/OTHER	0.3%	0.3%	96%

Banerji, OL et al. (0908.2033)

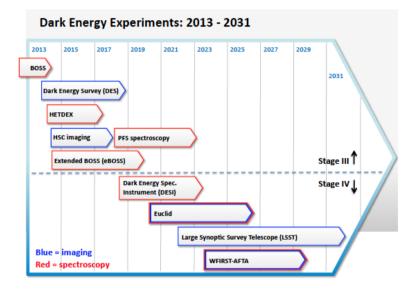
### Photometric Classification of Supernovae

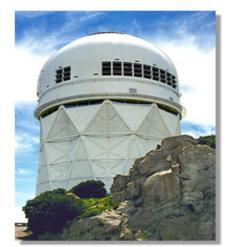


### Feature extraction with Wavelet

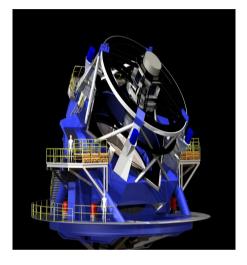


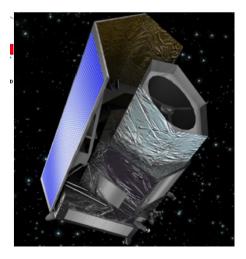
### The era of DESI, Euclid, LSST,...





Mayall 4-Meter Telescope







Telescope Field-of-View Survey Area

Camera Cadence

Raw Data Reduced

Catalog

### DES vs. LSST

DES 4 meters 3 sq-deg 5,000 sq-deg

570 megapixels 2 / yr / band <u>LSST</u> 8 meters 9.6 sq-deg 18,000 sq-deg

3,200 megapixels ~100 / yr / band

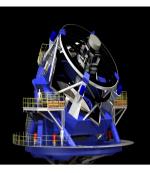
1 TB / night 2.5 PB

6 x 10<sup>8</sup> objects

2 x 10<sup>10</sup> objects

15 TB/ night

**Few 100 PB** 



### **Euclid Forecast**

What	Euclid	Before Euclid
Galaxies at 1 <z<3 estimates<="" good="" mass="" td="" with=""><td>~2x10<sup>8</sup></td><td>~5x10<sup>6</sup></td></z<3>	~2x10 <sup>8</sup>	~5x10 <sup>6</sup>
Massive galaxies (1 <z<3) <br="" w="">spectra</z<3)>	~few x 10 <sup>3</sup>	~few tens
Hα emitters/metal abundance in z~2-3	~4x10 <sup>7</sup> /10 <sup>4</sup>	~104/~102?
Galaxies in massive clusters at z>1	~2x10 <sup>4</sup>	~10³?
Type 2 AGN (0.7 <z<2)< td=""><td>~104</td><td>&lt;103</td></z<2)<>	~104	<103
Dwarf galaxies	~105	
T <sub>eff</sub> ~400K Y dwarfs	~few 10 <sup>2</sup>	<10
Strongly lensed galaxy-scale lenses	~300,000	~10-100
z > 8 QSOs	~30	None

### Big Data, Big collaborations: How many collaborators can one have?



2dF: 30 DES: 500 Planck:400 Euclid: 1200 LHC: 4000



Within the primates there is a general relationship between the size of the brain and the size of the social group. Scaling it to humans gives 150 people. This is the cognitive limit to the number of people with whom one can have stable interaction (Dunbar 1992). <sup>27</sup>

Some points for discussion (based on panel discussion of ATI-PS meeting in Jan 2016)

- What can we learn from other fields?
- What can we learn from other Big Data initiatives elsewhere in the world?
- How to connect applications to computer science foundations (networks etc.)
- How to move from 'deterministic' algorithms to computation in the presence of uncertainties.
- Would Big Data lead to more (or less) 'deep thinking' in Physics problems?

# Summary

- Science is going 'industrial revolution'
- Both spatial and time domains
- Great training of PhDs, beyond academia.
- Will Big Data produce bigger knowledge?
  - (it depends it part on Nature...)